



ISTITUTO NAZIONALE
DI GEOFISICA E VULCANOLOGIA



This presentation on the Trans-National Access (TNA) project titled

Validating the Swarm S4 index over Africa using the Eswua Database (VSS4AED)

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TNA project, INGV Node

Participants:

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SWMC team: Ola Abu Elzz and Ahmed Yassen (EJUST)

Outline

- 1 Abstract**
- 2 Introduction and Objectives**
- 3 Model Validation**
- 4 Measured data and Model parameters**
- 5 Comparing with the manual scaled ionograms.**
- 6 Conclusion**
- 7 Research outcome:
Publications**
- 8 The Role of PITHIA-NRF in My Project**

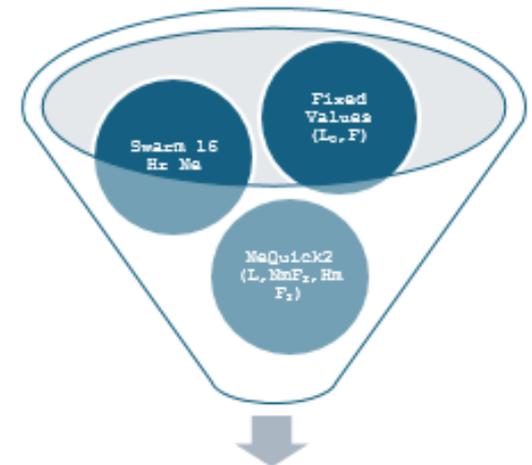
Abstract

- The **amplitude scintillation index S₄** is a measure of the intensity fluctuations of Global Navigation Satellite System (GNSS) signals caused by ionospheric irregularities.
- Recently, **The INGV** proposed and designed **a model to estimate S₄** from the 16 Hz density data measured by the Swarm satellites.
- The model based **on Rino's theory** of weak scattering and **the NeQuick2 model** of the electron density profile.
- They used the **Swarm data** to derive the **spectral slope** and **the variance of the electron density** at the peak of the irregularity layer, which are the key parameters for the scintillation model.
- Then they used the NeQuick2 model to reconstruct the irregularity layer.
- In this project, my contribution was to validate the INGV proposed model using Malindi observatory data that intersect with Swarm trajectories.

The INGV proposed model can provide global information on scintillation, especially in areas where ground-based GNSS receivers are unavailable.

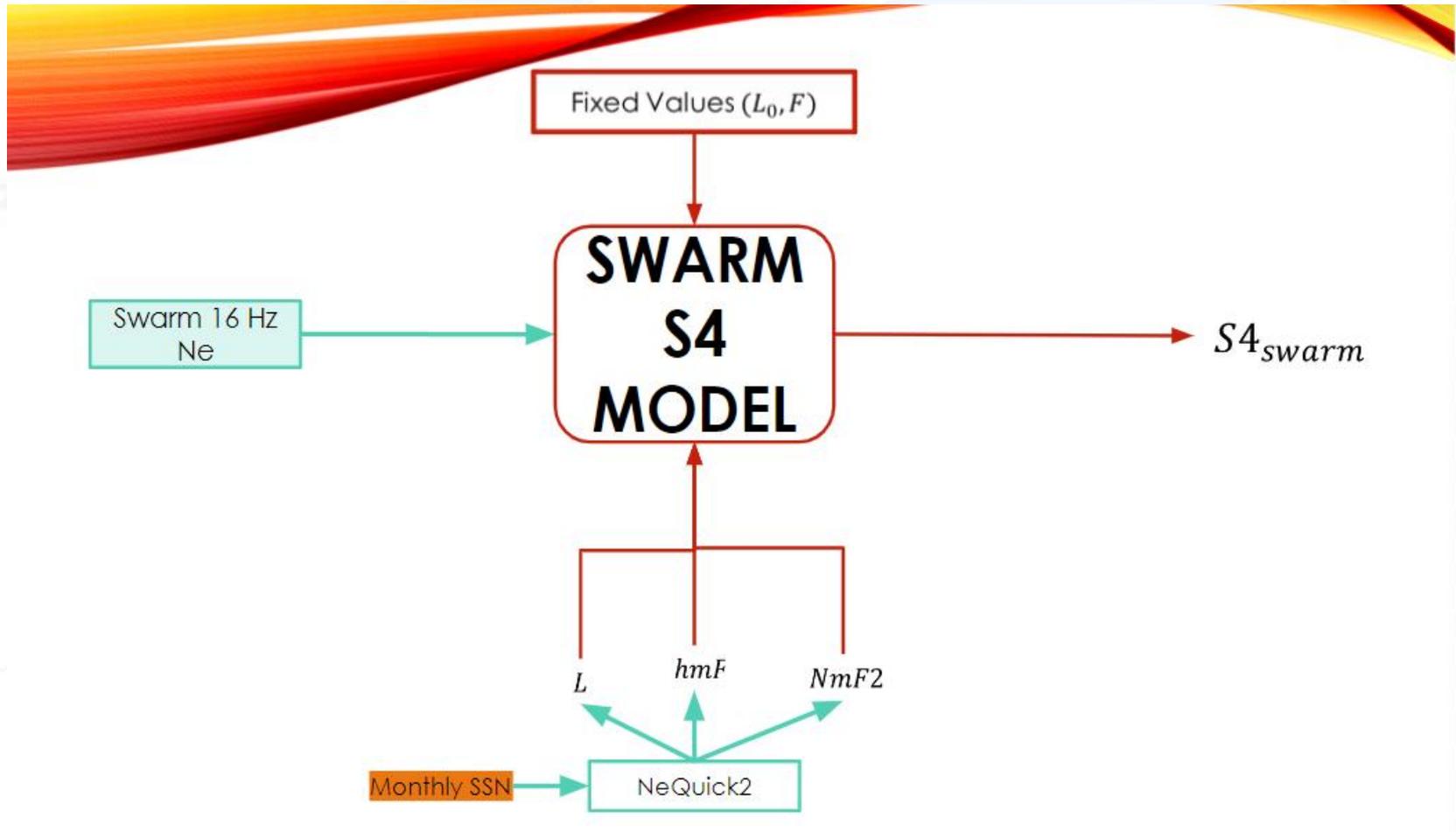
The WAM scintillation model approach for high latitudes and for low latitudes procedure:

- Assume virtual GNSS satellites and ground GNSS receivers that are coincident with Swarm trajectories.
- Compute the profile of the electron density from the virtual GNSS receivers on the ground to the GNSS satellites at 20,000 km over and below the Swarm path and altitude.
- The NeQuick electron density profile is related to the ground-based Total Electron Content (TEC) measurements from the Madrigal database.
- The importance of this choice for reliable results, used for the validation.
- use the same ray path that the ground-based GNSS scintillation receiver observed.



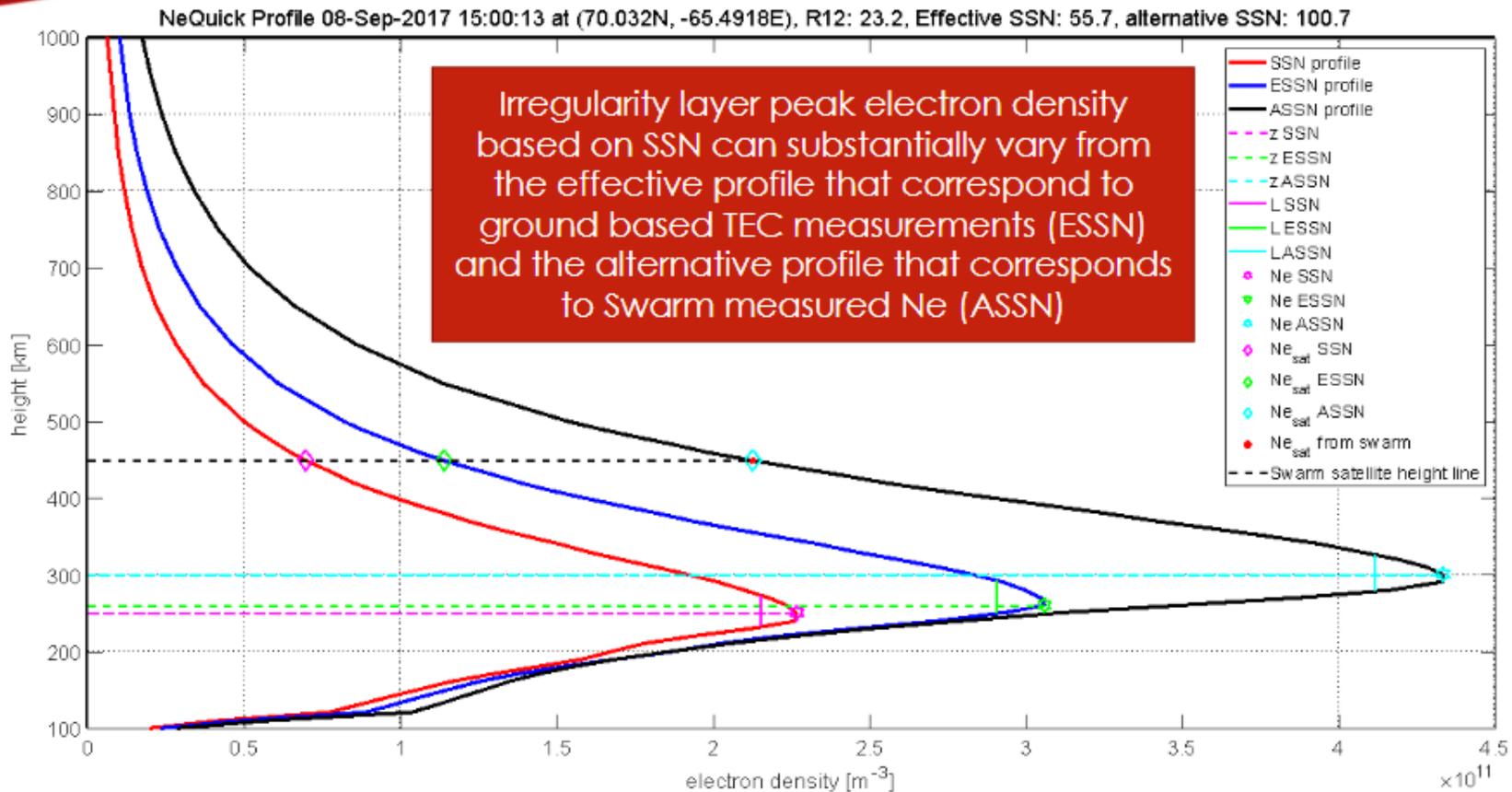
INGV proposed model to derive Swarm S₄

INGV Model's parameters

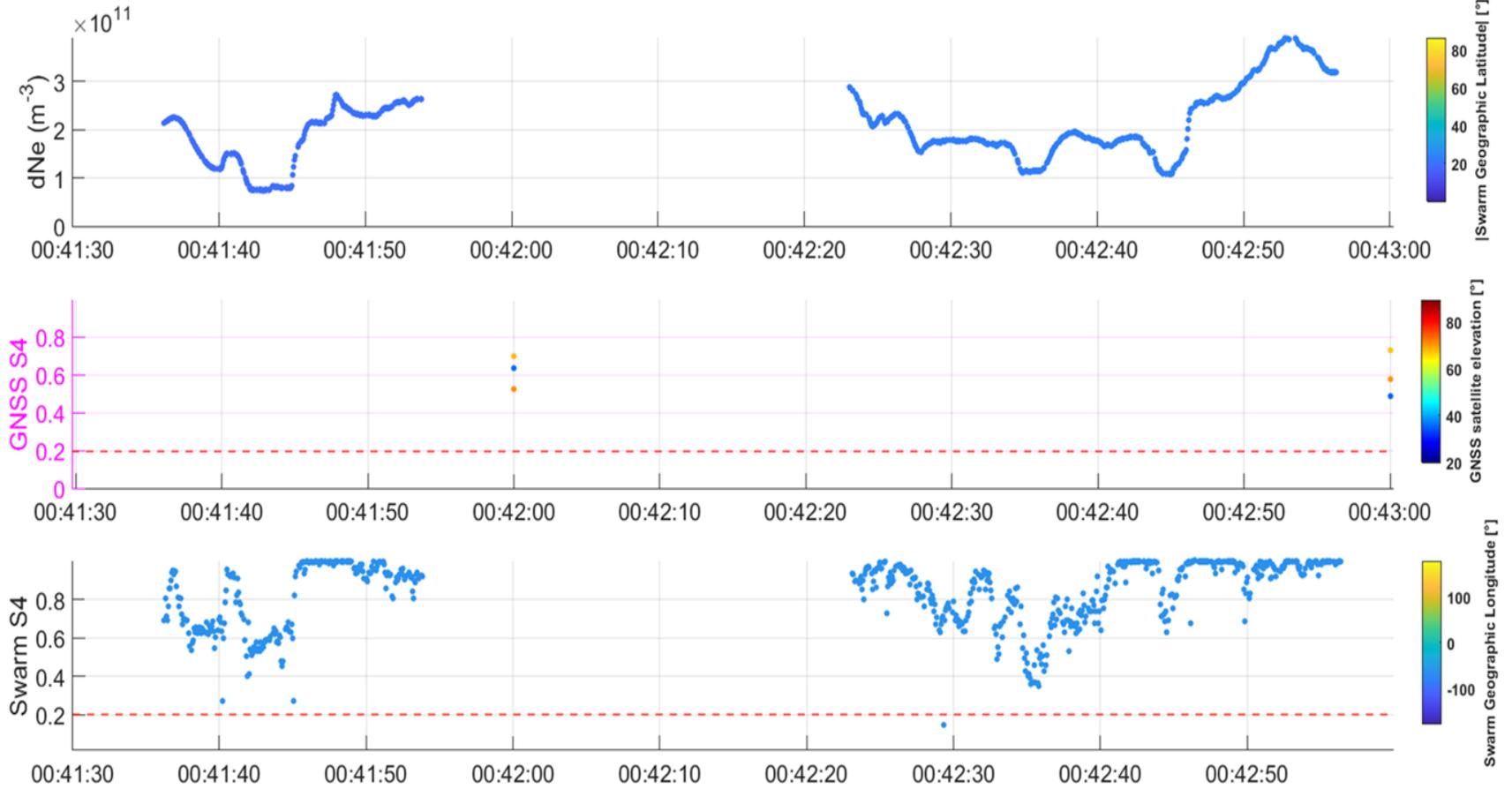


The Three NeQ-Ne profiles models

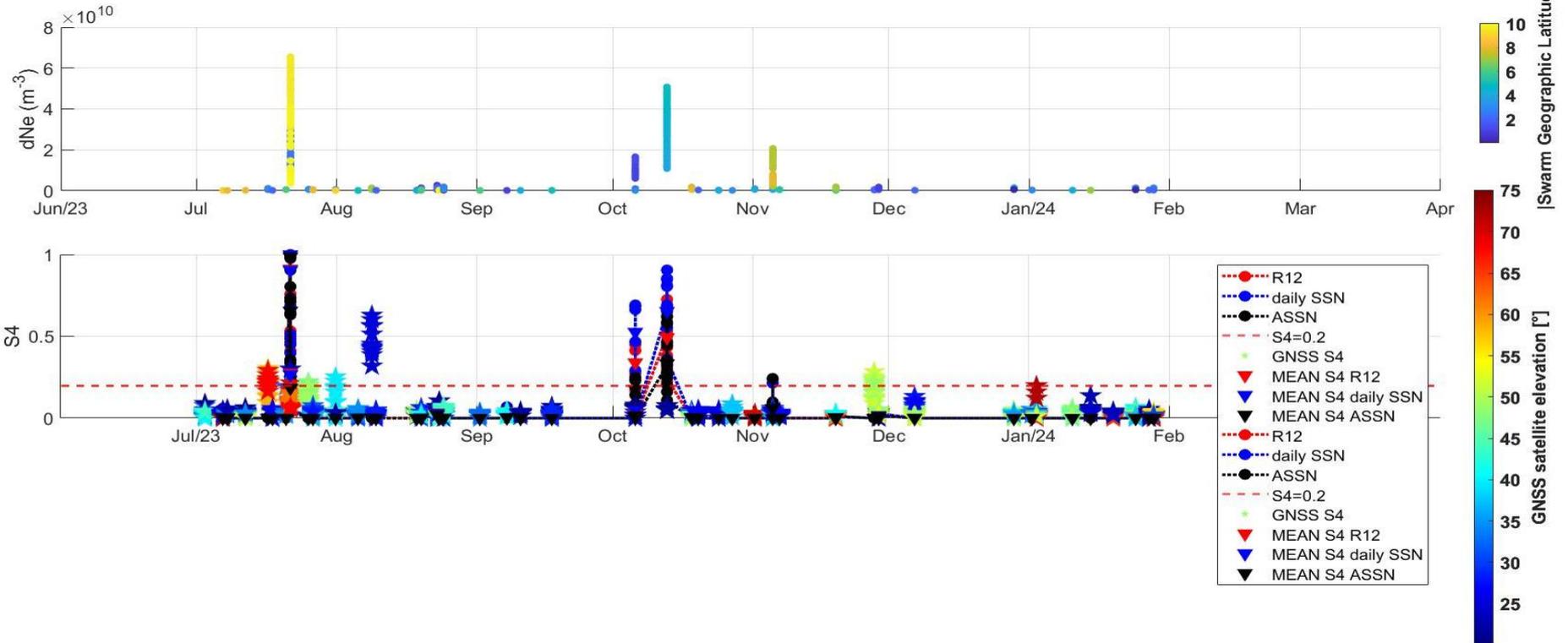
EXAMPLE OF NEQUICK2 PROFILES



SwarmA - 02-Sep-2021 20:47 --> 13-Apr-2023 17:02 -- L0=500 [m]



SwarmA - 06-Jul-2023 20:10 --> 28-Jan-2024 13:41 -- L0=500 [m] || $L_{NeQuick}$ estimated from the profile using the TEC/NmF2 equation



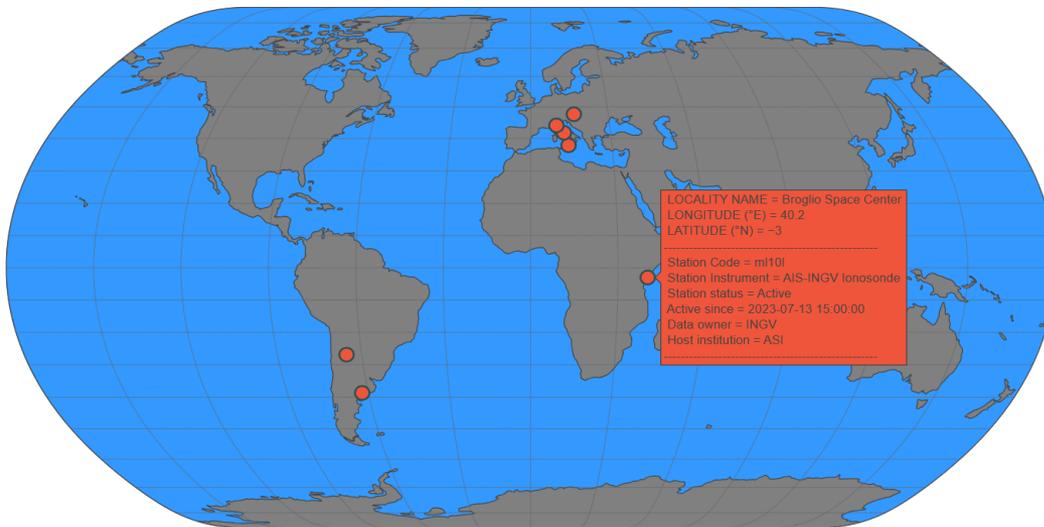
Malindi Ionosonde

- The Swarm S4 index using GNSS data collected from Malindi station in Africa, where ionospheric irregularities are more pronounced due to geomagnetic equatorial effects.

Ionosonde Station	Latitude	Longitude
Malindi	-3°N	40.2° E



INGV IONOSPHERIC MONITORING NETWORK



● GNSS receivers for TEC and scintillation monitoring ● Ionosondes



The New Observatory for Real-time Ionospheric Sounding over Kenya-
NORISK project (<http://norisk.rm.ingv.it/>)

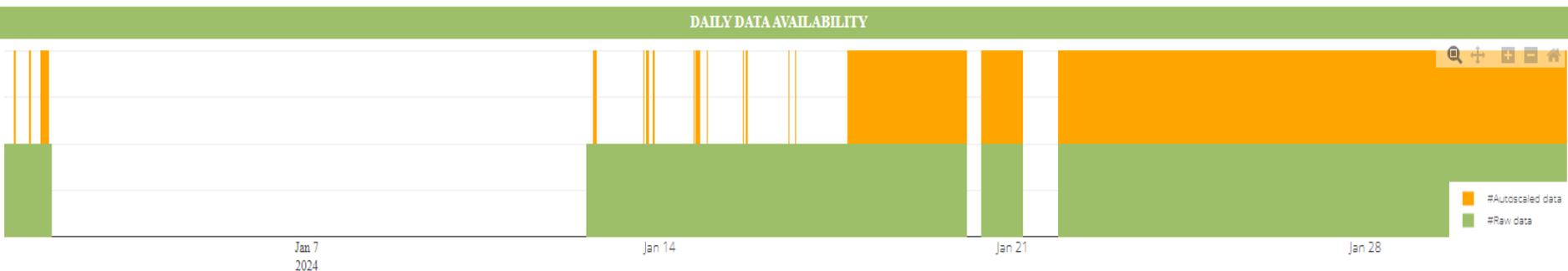
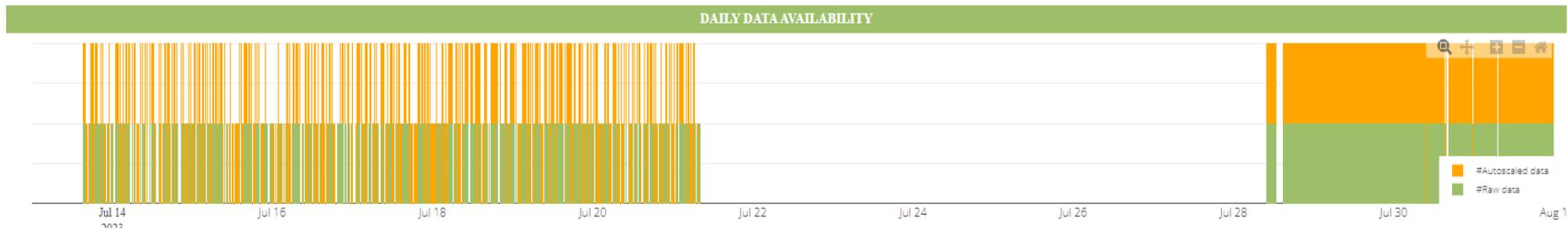
Data availability

- Malindi Ionosonde data started on 13/7/2023 up to now.

In the validation process:

- We started our analysis by using the Autoscaled ionograms.
- The data range is from 5/8/2023 to 28/01/2024.

<http://www.eswua.ingv.it/>





SELECT STATION:

AFRICA - ml101

Station info and data availability

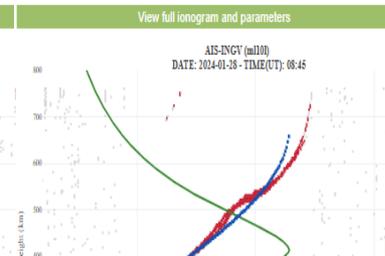
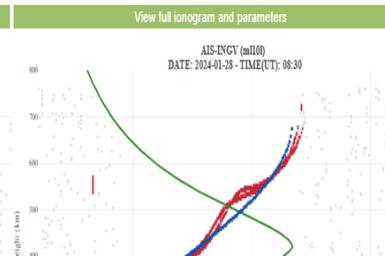
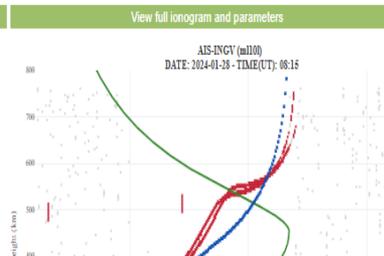
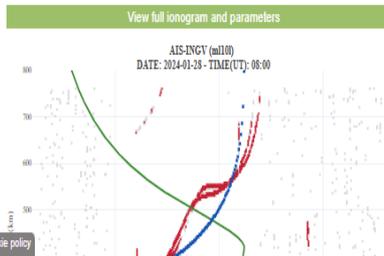
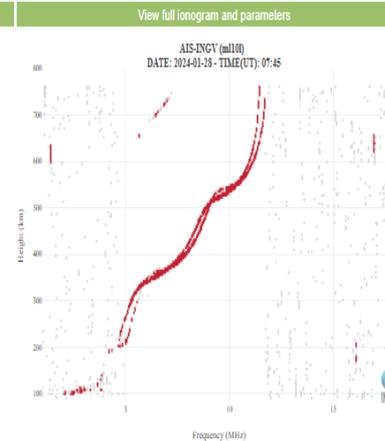
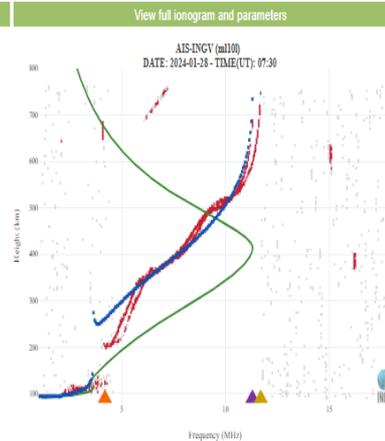
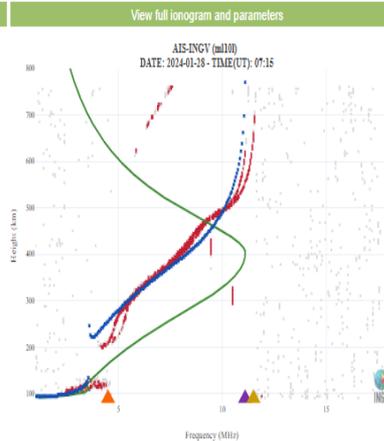
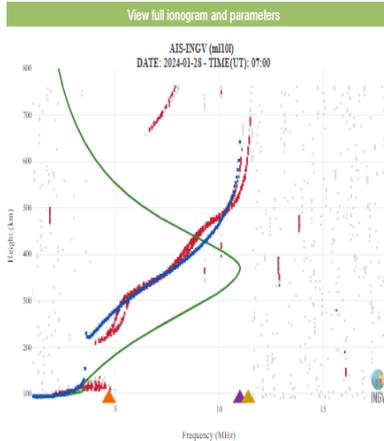
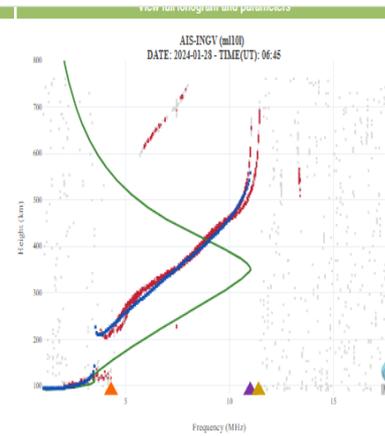
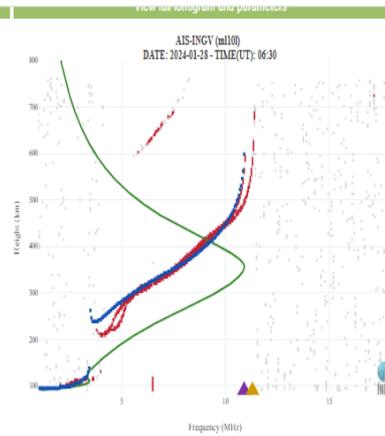
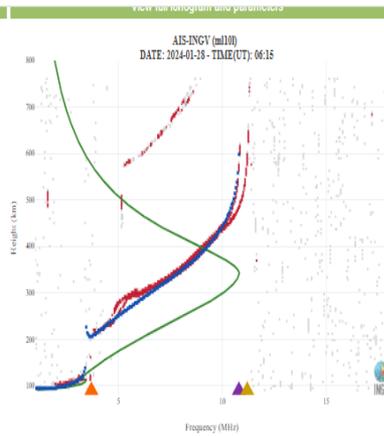
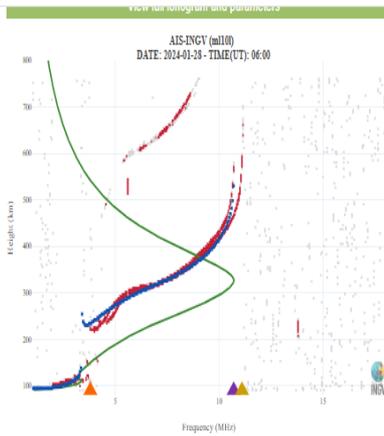
TIME RANGE (UTC):

2024/01/28 00:00 - 2024/01/28 12:00

FILTER SPECIFIC INSTANT:

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APPLY FILTERS



Model validation

Methodology

- 1- **Finding the time Synchronization (matching)** between the 3 different instruments.
- 2- **Compute** the **NmF2**, **hmF2** from the Ionosonde (Malindi station).
- 3- **Compute** slap thickness (**L**) from the **NmF2** (ionosonde) and **TEC** from GPS measurements.
- 4- **Calculating** the **mean and standard deviation** values from each parameter (**NmF2**, **hmF2** and **L**).

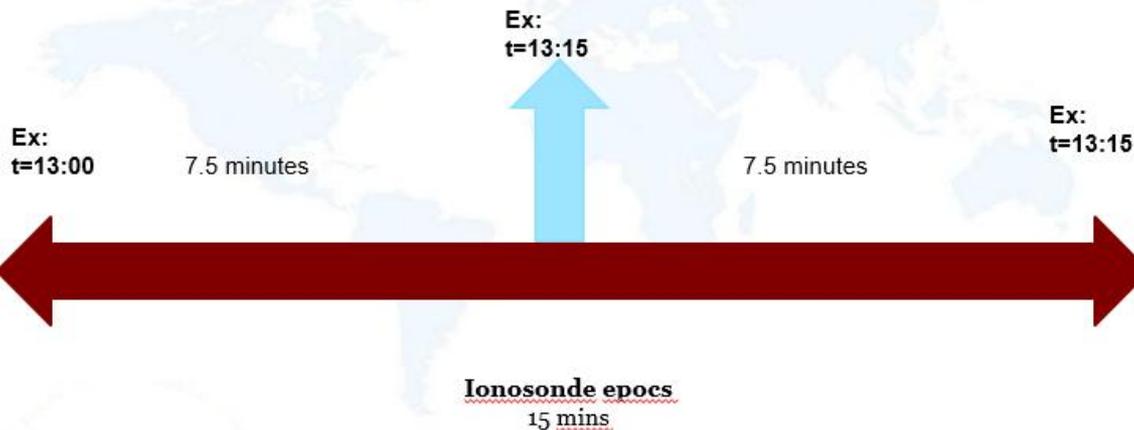
Time Matching

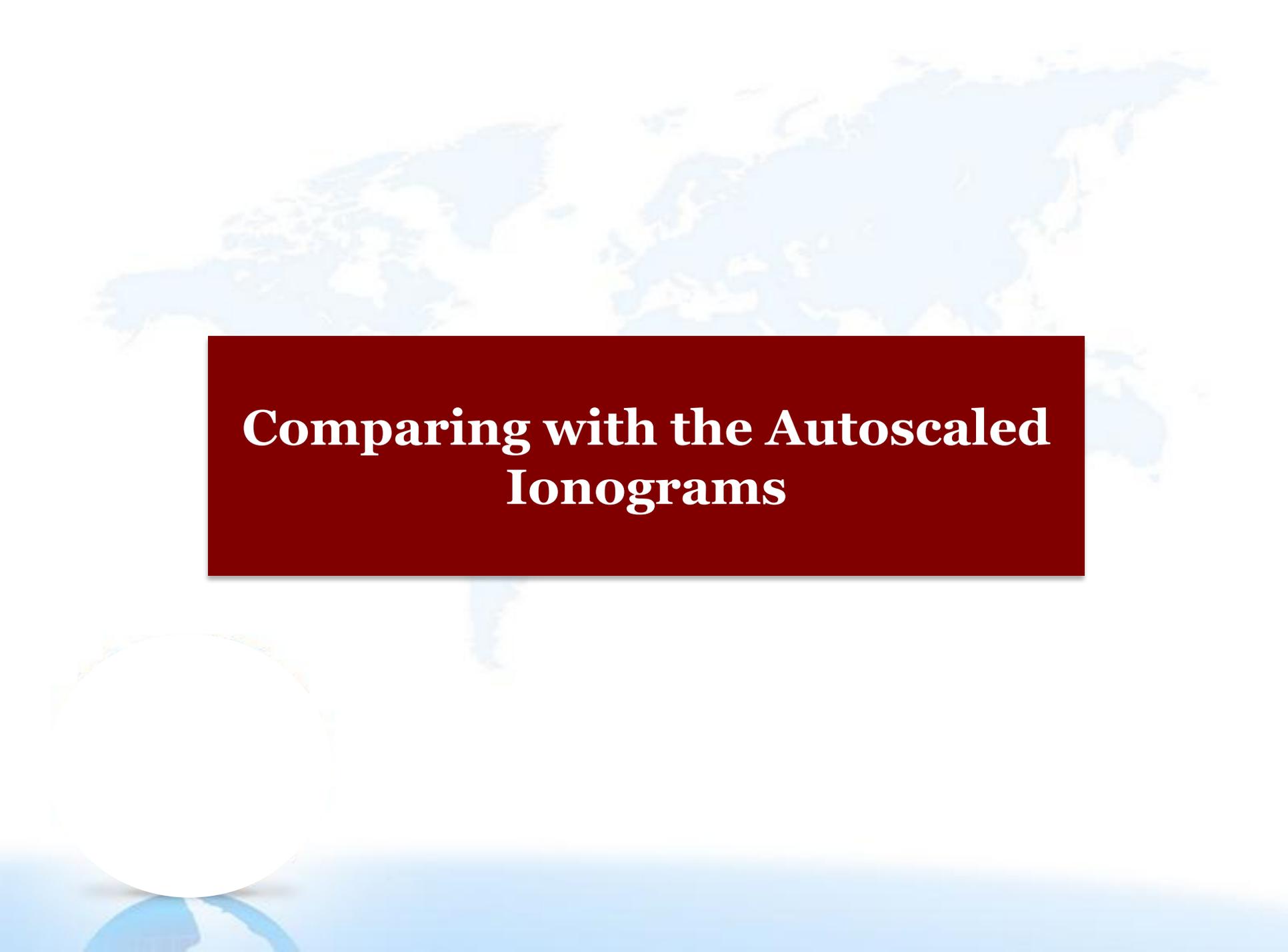
Swarm Time
measurements
Every 16 Hz

Ionosonde epocs
Every
15 mins

GPS Time
measurements
Every 1 min

The reference time is the ionosonde epoch.

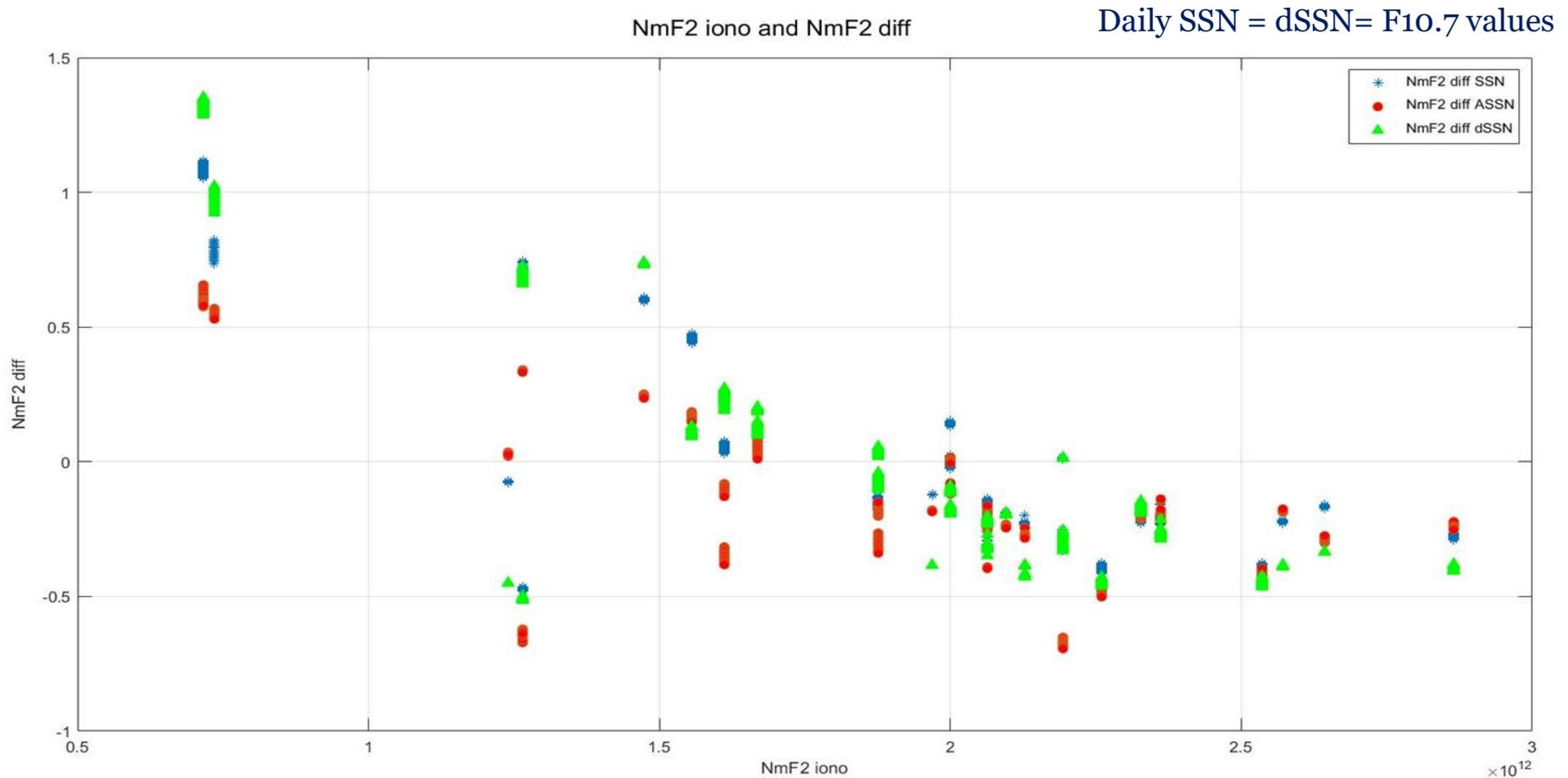




Comparing with the Autoscaled Ionograms

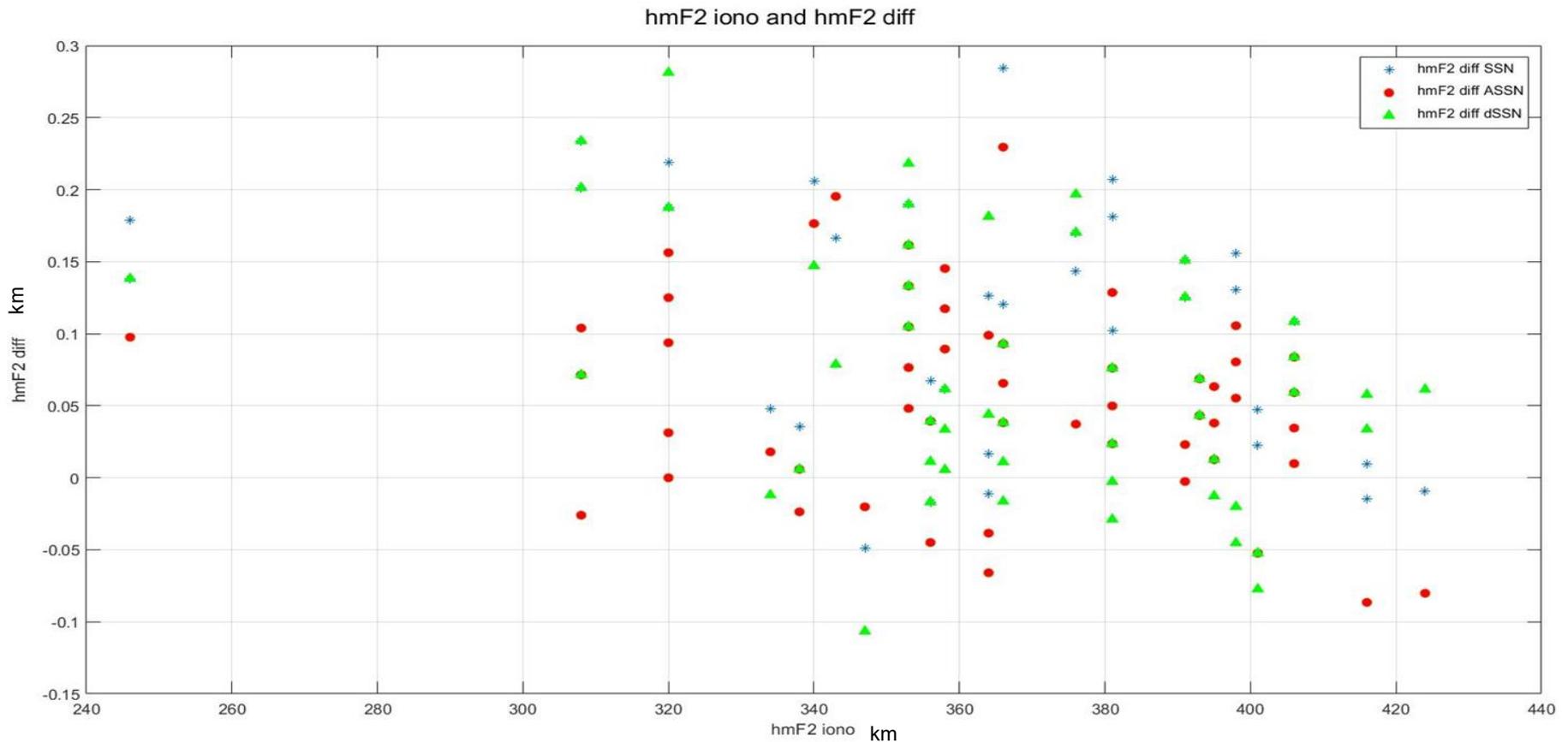
NmF2 Validation

➤ The distribution of the modelled NmF2 as calculated from 3 NeQ models with the NmF2 as observed from Malindi ionosonde.



hmF2 Validation

- The distribution of the modelled hmF2 as calculated from 3 NeQ models with the hmF2 as observed from Malindi ionosonde.



Slab thickness comparison

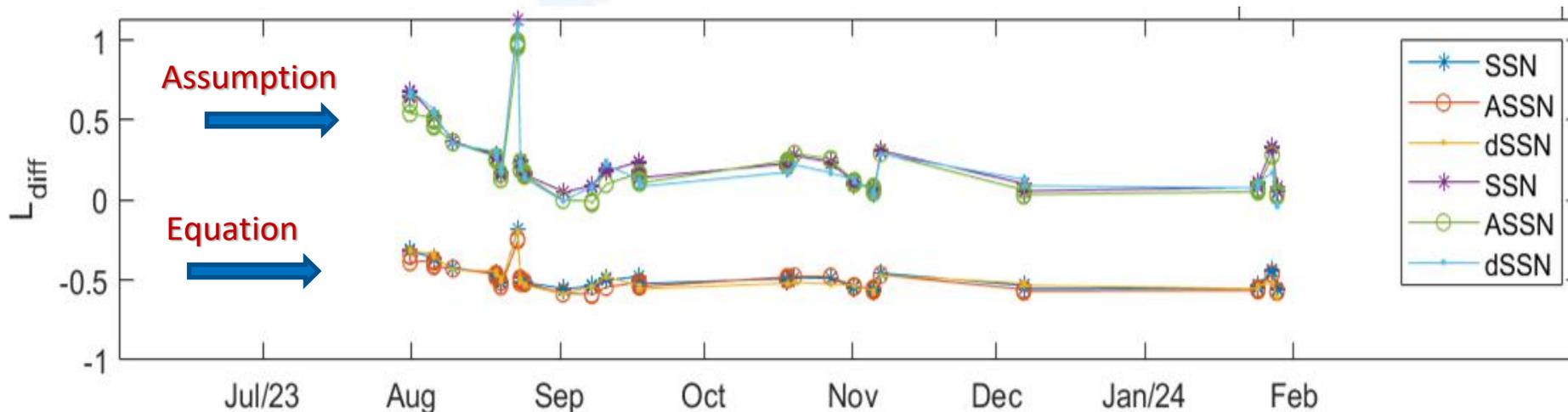
➤ The distribution of the modelled L as calculated from 3 NeQ models with the L as computed from $L_{\text{assumption}}$ and L_{equation} (TEC calculation).

Assumption 75%

$$L = NmF2 * 75\%$$

TEC_calculation

$$L_{\text{eq}} = \text{TEC} / NmF2$$

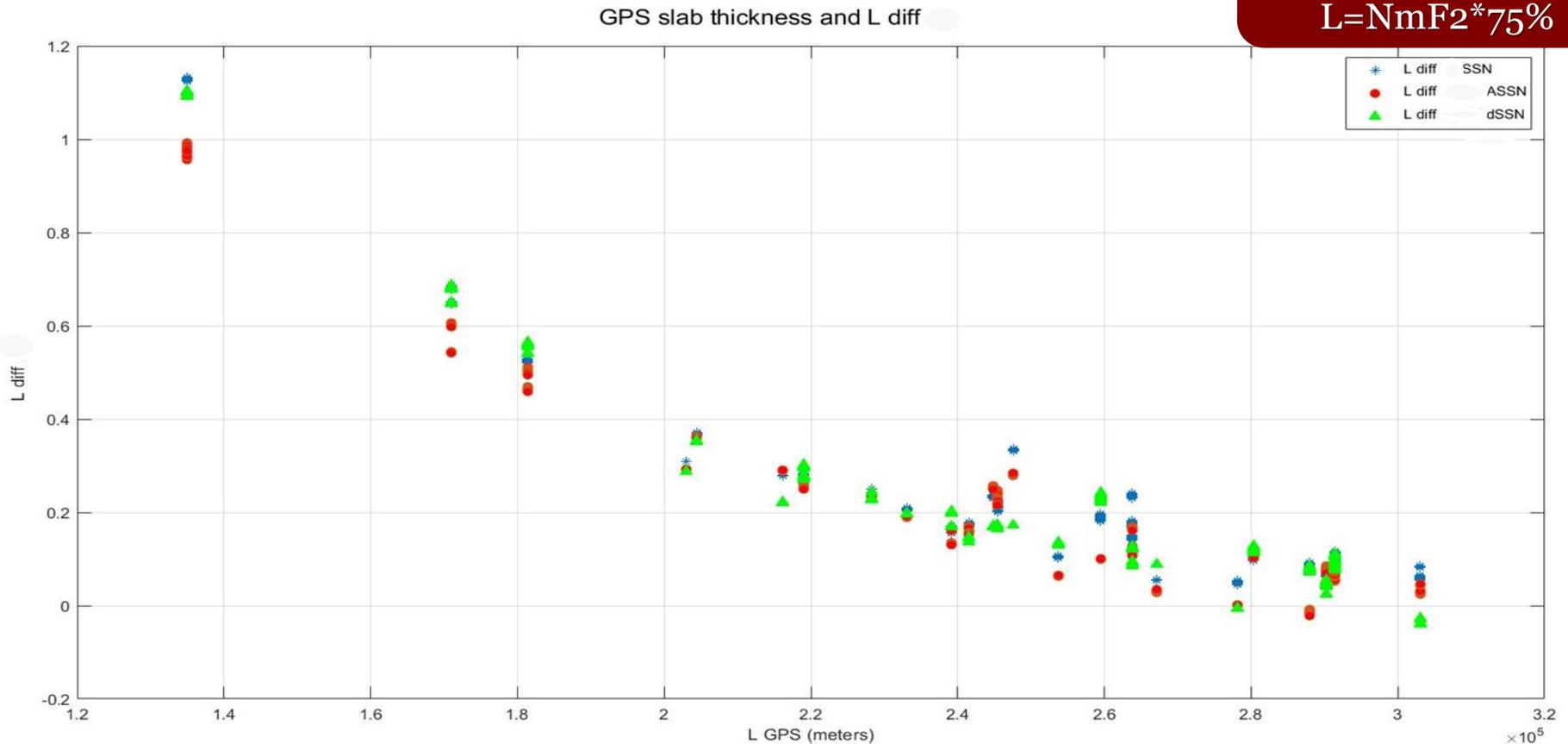


Slab thickness comparison

➤ The distribution of the modelled L_{diff} (assumption) as calculated from the 3 NeQ models with the L_{GPS} as observed from Malindi ionosonde.

Assumption 75%

$$L = NmF2 * 75\%$$

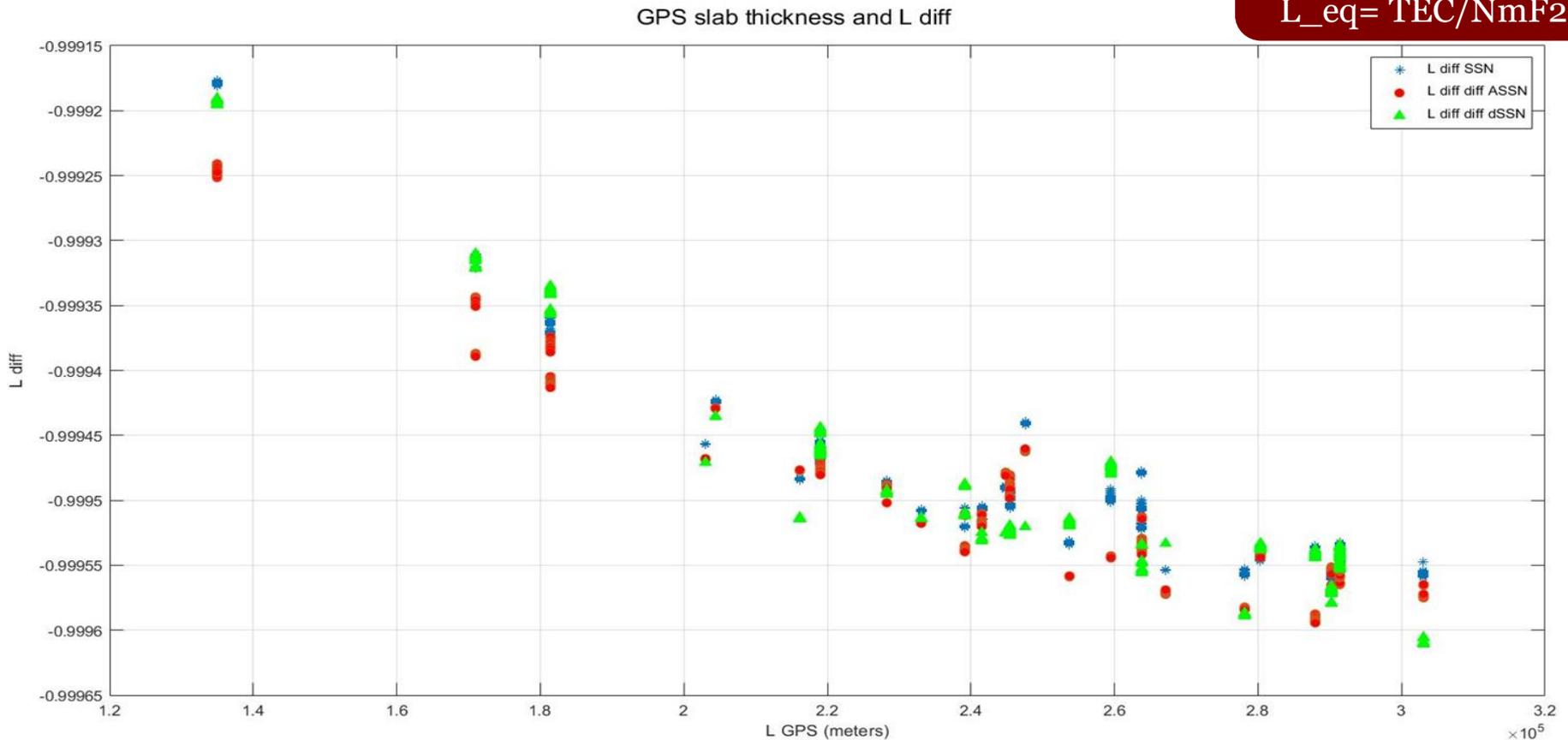


Slab thickness comparison

➤ The distribution of the modelled L_diff_eq (TEC/NmF2) as calculated from 3 NeQ models with the L_GPS as observed from Malindi ionosonde.

TEC_calculation

$L_eq = TEC / NmF2$



Data comparing

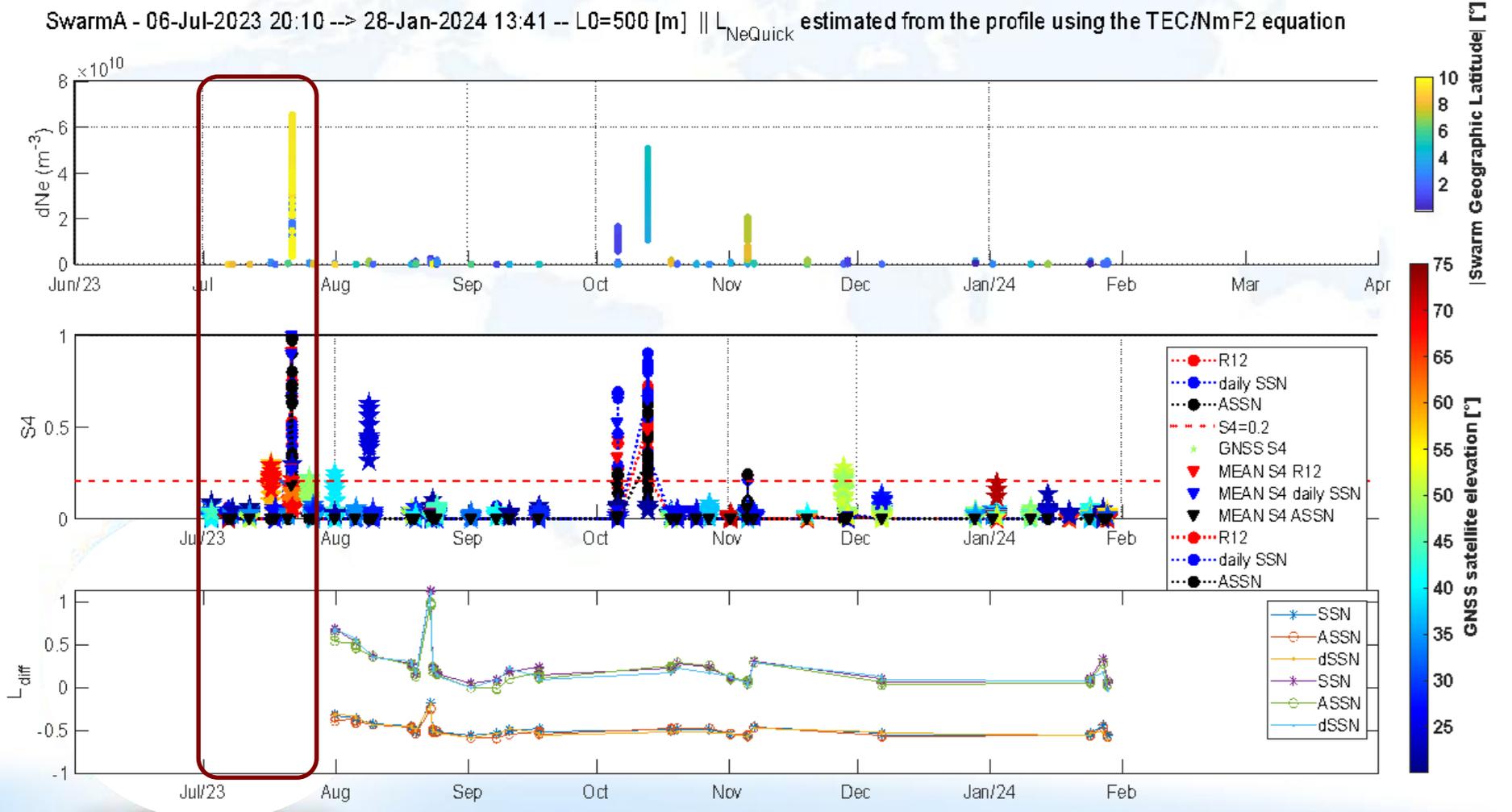
Three input model parameters (L,NmF2 and hmF2):

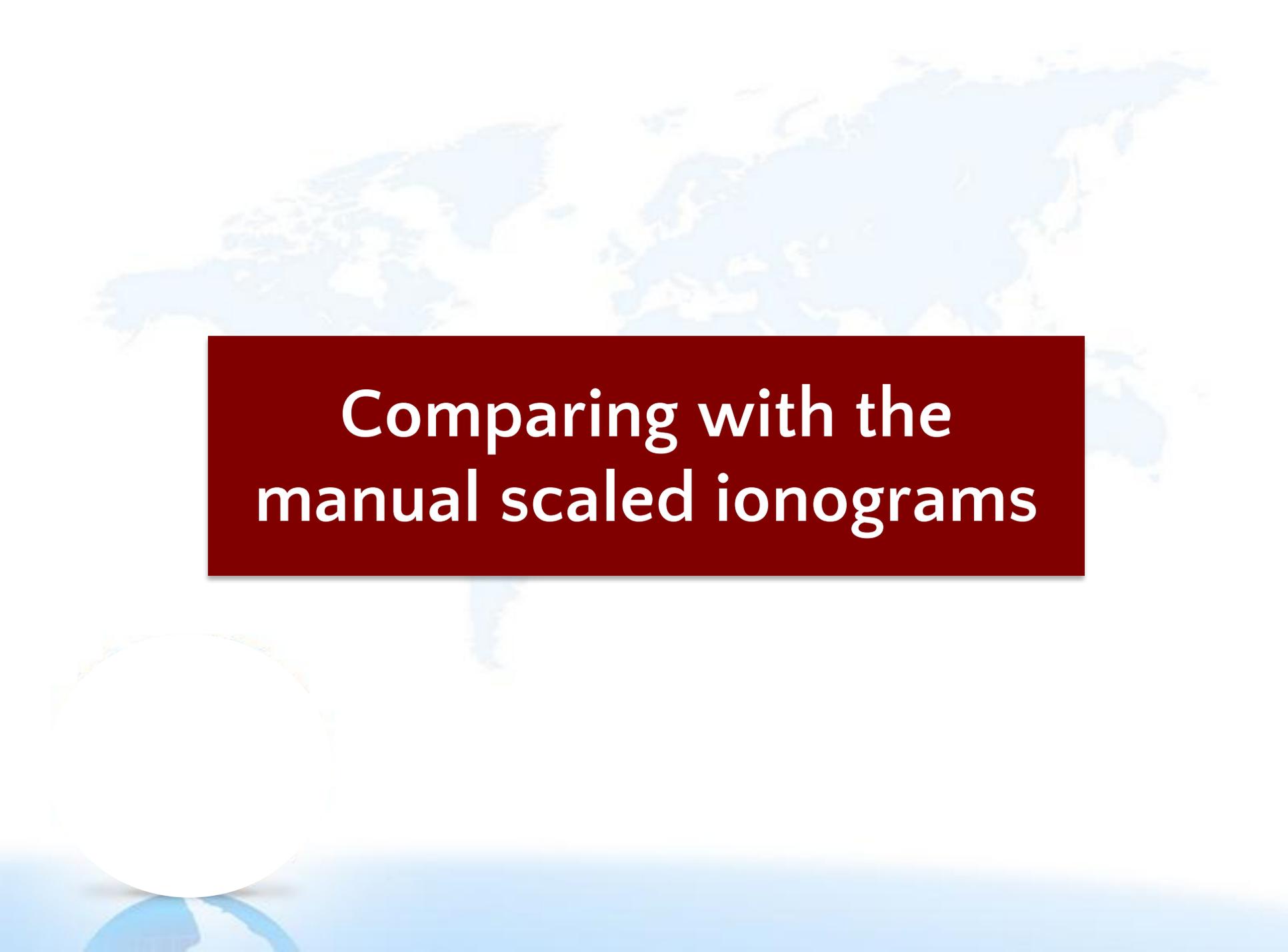
➤ Data comparing: the Auto scale data

Parameter	diff_SSN_mean	diff_ASSN_mean	diff_dSSN_mean
L_diff	-0.49707	-0.49707	-0.49178
L_diff_eq	0.2538	0.22242	0.23166
NmF2_diff	-0.019704	-0.13407	-0.06175
hmF2_diff	0.096985	0.057585	0.06853

Parameter	diff_SSN_std	diff_ASSN_std	diff_dSSN_std
L_diff	0.0913	0.0802	0.0960
L_diff_eq	0.2480	0.2224	0.2533
NmF2_diff	0.3696	0.2849	0.4486
hmF2_diff	0.0706	0.0729	0.0816

Why we are thinking to include the manual scale data ?

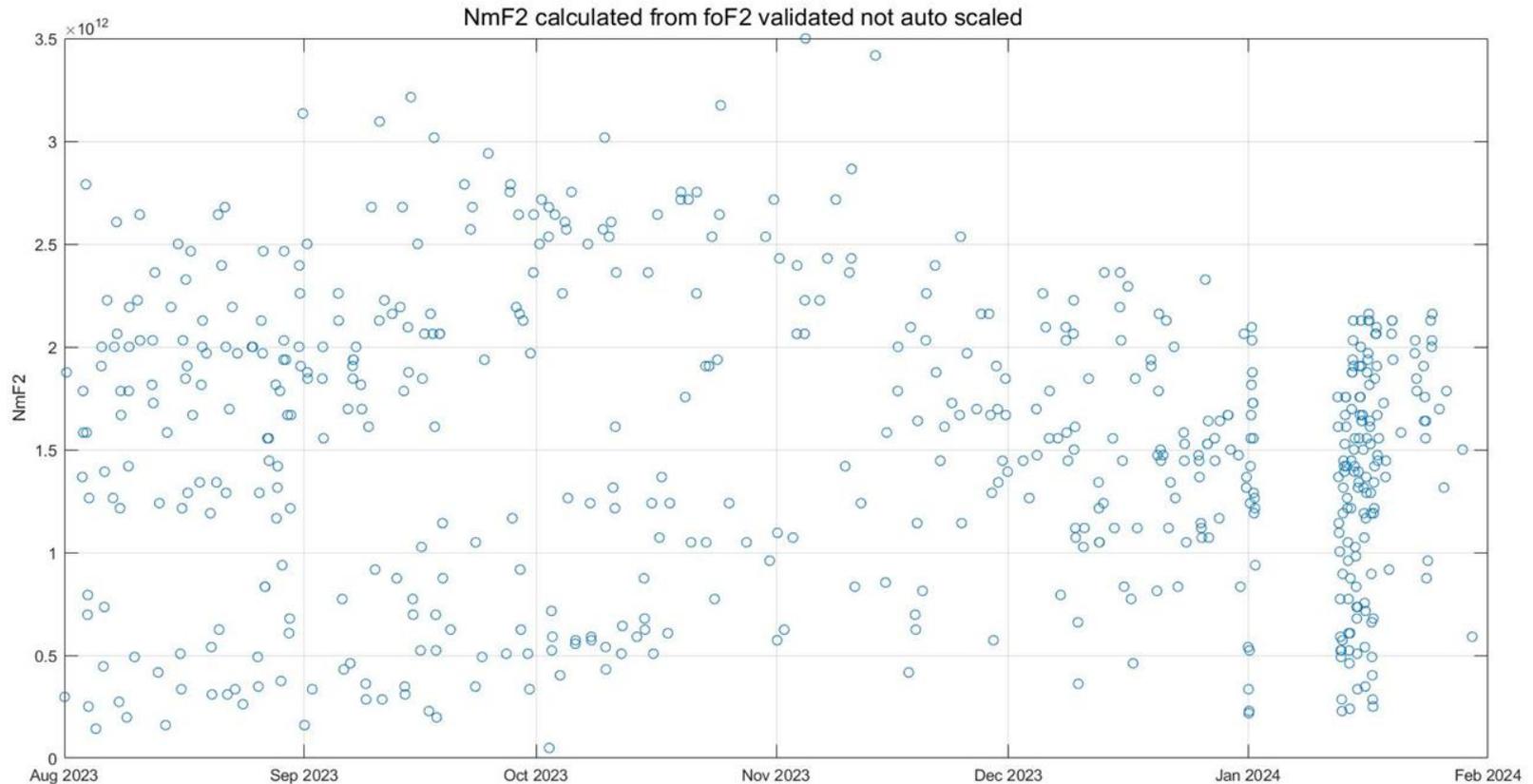


A faded world map in light blue and white tones serves as the background for the slide. In the bottom left corner, there is a small, stylized globe showing blue oceans and white landmasses.

Comparing with the manual scaled ionograms

Manual scaled ionogram

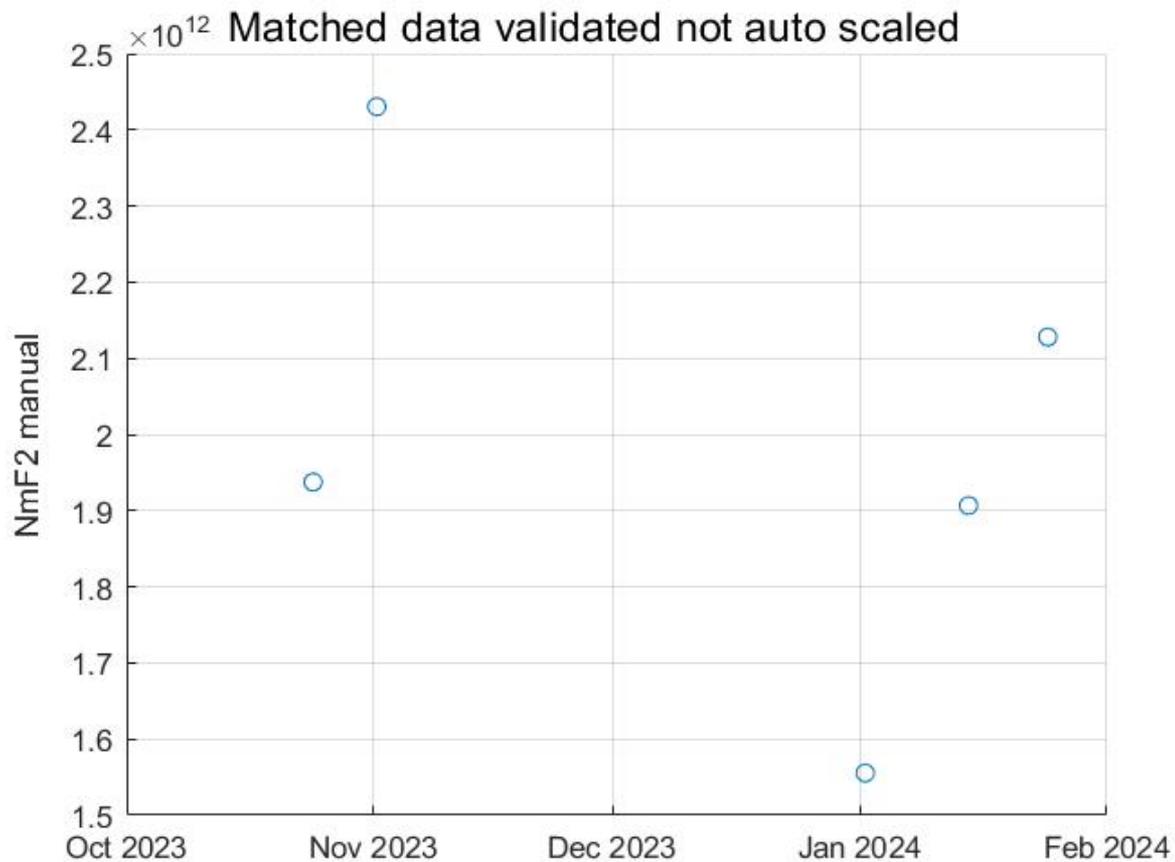
- **NmF2** as calculated from the validated ionograms, but not autoscaled.



Number of manual scaled ionogram= 566

Manual ionogram data

- Matching manual scaled ionograms with Swarm measurements



Number of matched data
between ionosonde with
swarm = 5

Conclusion

- For the Auto scaled ionograms:
 - ✓ The **diff_ASSN_mean** of three NeQ parameters (NmF2_ASSN, hmF2_ASSN, and L_ASSN) give the **smallest** values.
 - ✓ The **L_diff_SSN** values are similar to the **L_diff_ASSN** values.
 - ✓ The **diff_ASSN_std** for the three NeQ parameters is the **smallest** value except for the (hmF2).
- For the Manual scaled ionograms:
 - ✓ Due to the few matched ionograms with the swarm trajectories, we couldn't compute the statistical analysis on the data.

Research outcome

Publications:

1. Validating the amplitude scintillation index from Swarm satellites faceplate measurements against the GNSS scintillation receiver measurements in Malindi, Conference paper, SIF National Congress 2024, Rome, Italy, 9 - 13, September 2024.
1. Validating Swarm S4 index from the faceplate electron density measurements over a pair of collocated ionosonde and GNSS receivers in Sharjah, UAE, and Malindi, Kenya. Conference paper, Conference of 4th URSI Atlantic Radio Science Meeting, AT-RASC 2024, Gran Canaria, Spain, 19-24 May 2024.
DOI: [10.46620/URSIATRASC24/PFRO1041](https://doi.org/10.46620/URSIATRASC24/PFRO1041)

The Role of PITHIA-NRF in My Project

The Trans-National Access (TNA) is one of the objectives of the PITHIA-NRF project.

- I have attended two TNA calls:
 - ✓ in 2024 at the INGV node (project: VSS4AED) and
 - ✓ in 2025 at the NOA node (project: CATNPRO).
- PITHIA-NRF infrastructure was important in the successful accomplishment of my project, primarily by providing access to high-quality data, computational resources, and expertise that were otherwise difficult to access.
- The collaboration of the PITHIA-NRF TNA programme enhanced communication and the exchange of ideas and experience between researchers from different institutes.

Acknowledgement

- I acknowledge the research infrastructures and the access provider TNA Node of INGV of the PITHIA-NRF project (<https://www.pithia-nrf.eu/>). The PITHIA-NRF project has received funding from the European Union's Horizon 2020 research and innovation programme under grant agreement No 101007599.